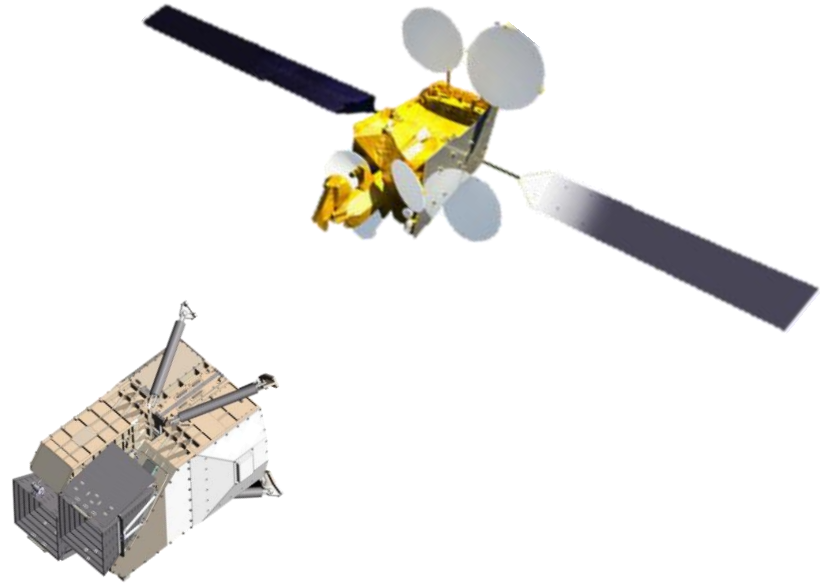
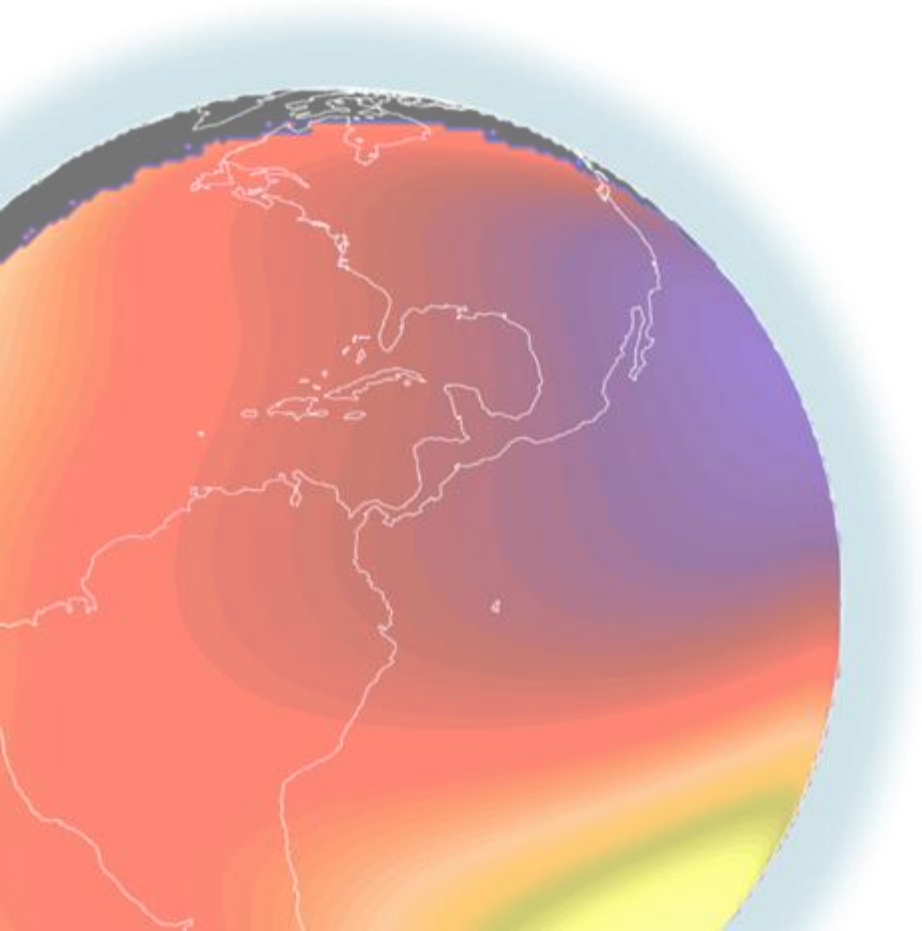


# **Imaging the Boundary Between Earth and Space – A Preview of Space Weather Data from the Global-scale Observations of the Limb and Disk (GOLD) Mission**



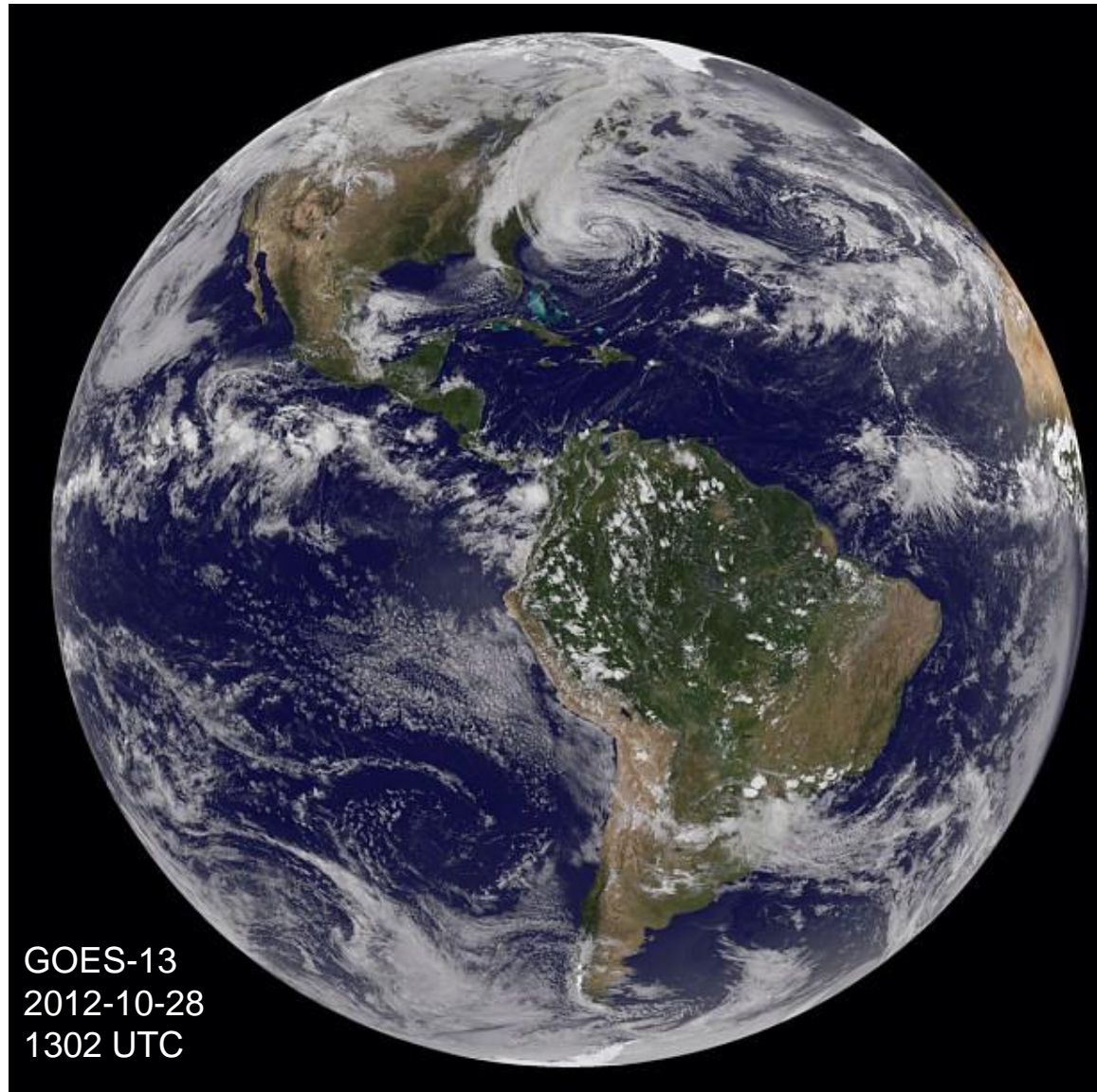
**Richard Eastes (UCF/FSI),  
William McClintock (CU/LASP),  
Alan Burns (NCAR) and  
the GOLD Science Team**



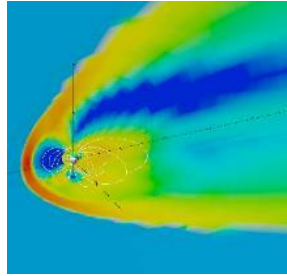
***GOLD will make unprecedented images of neutral temperature and composition in the upper atmosphere's***

**GOLD images the disk and limb from geostationary orbit**

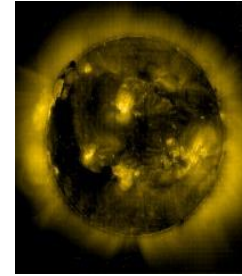
**Full disk images at 30-minute cadence**



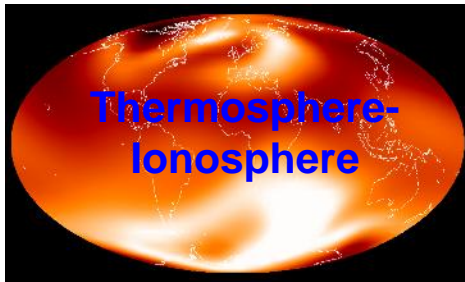
## *Forcing from Above*



**Science Question 1 (Q1).**  
How do geomagnetic storms alter the temperature and composition structure of the thermosphere?

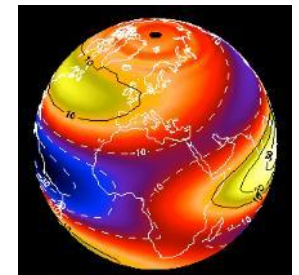


**Q2.** What is the global-scale response of the thermosphere to solar extreme-ultraviolet variability?



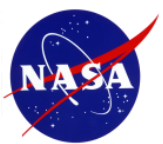
**Q4.** How does the nighttime equatorial ionosphere influence the formation and evolution of equatorial plasma density irregularities?

**Q3.** How significant are the effects of atmospheric waves and tides propagating from below on thermospheric temperature structure?



## *Forcing from Below*





# Ultraviolet Imaging from Geostationary Orbit

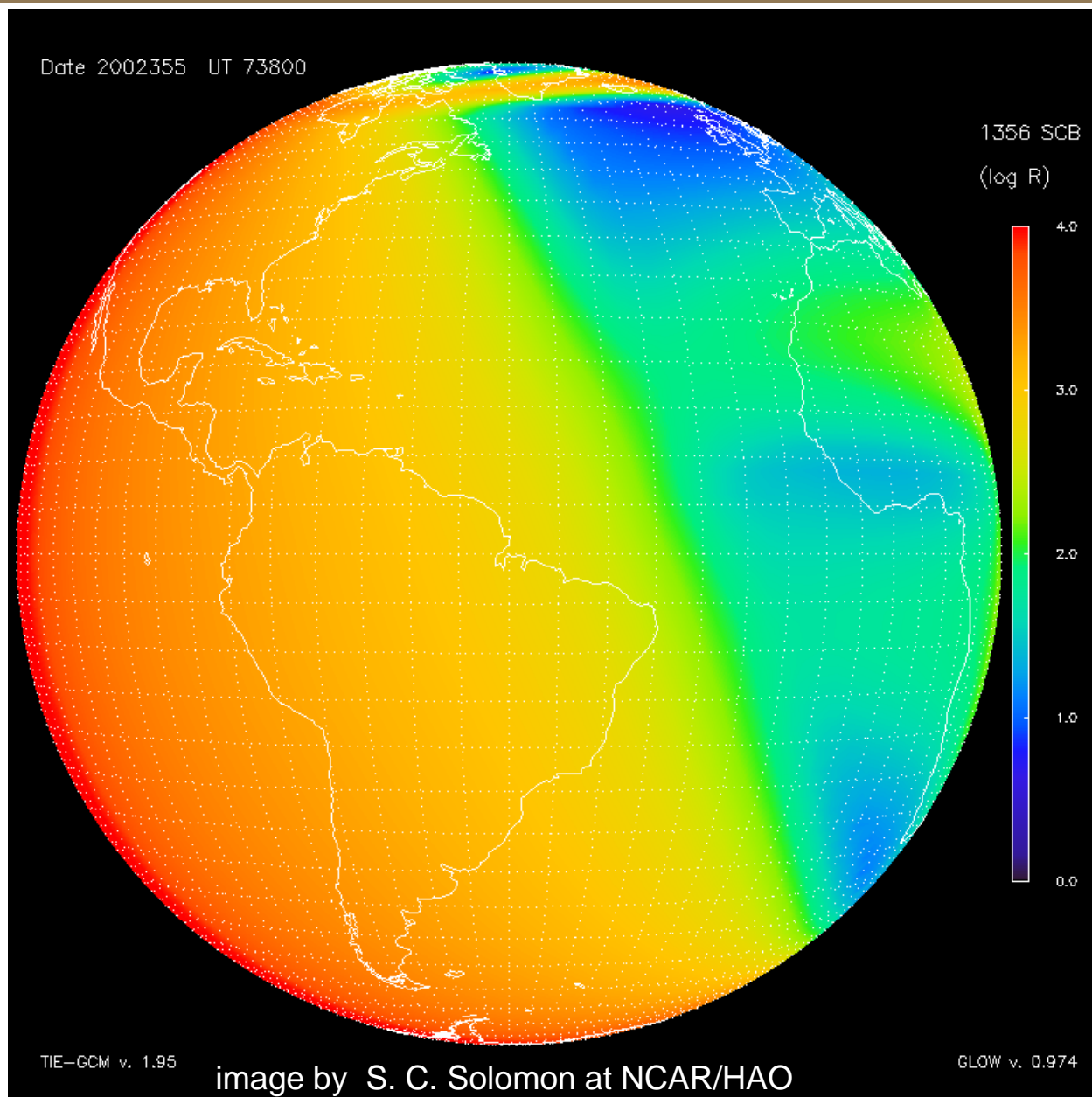
**GOLD**

**Simulated GOLD  
image of oxygen  
(135.6 nm) emissions**

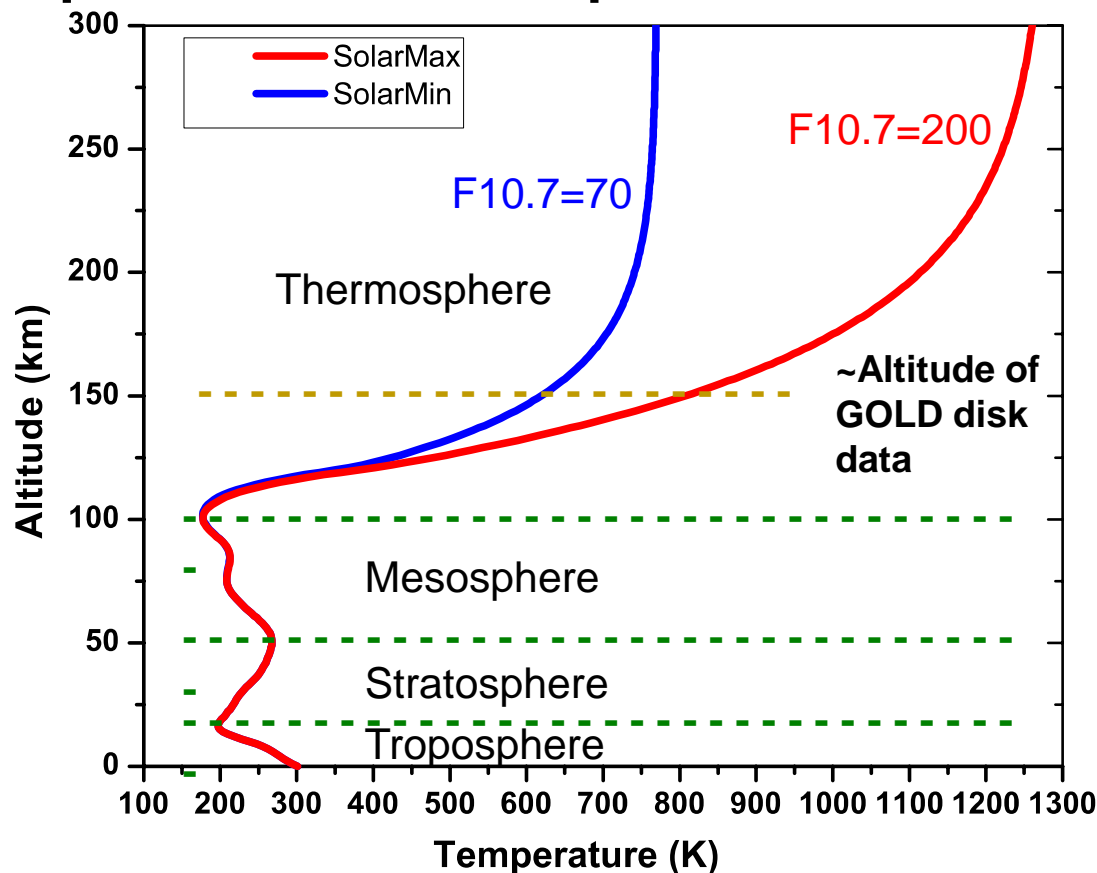
**Simultaneously  
images N<sub>2</sub> emissions  
on dayside**

**Emissions provide  
key data for bubbles,  
satellite drag, and  
electron densities**

**Provides data to  
advance predictions  
of assimilation  
models and of  
geomagnetic storm  
effects**

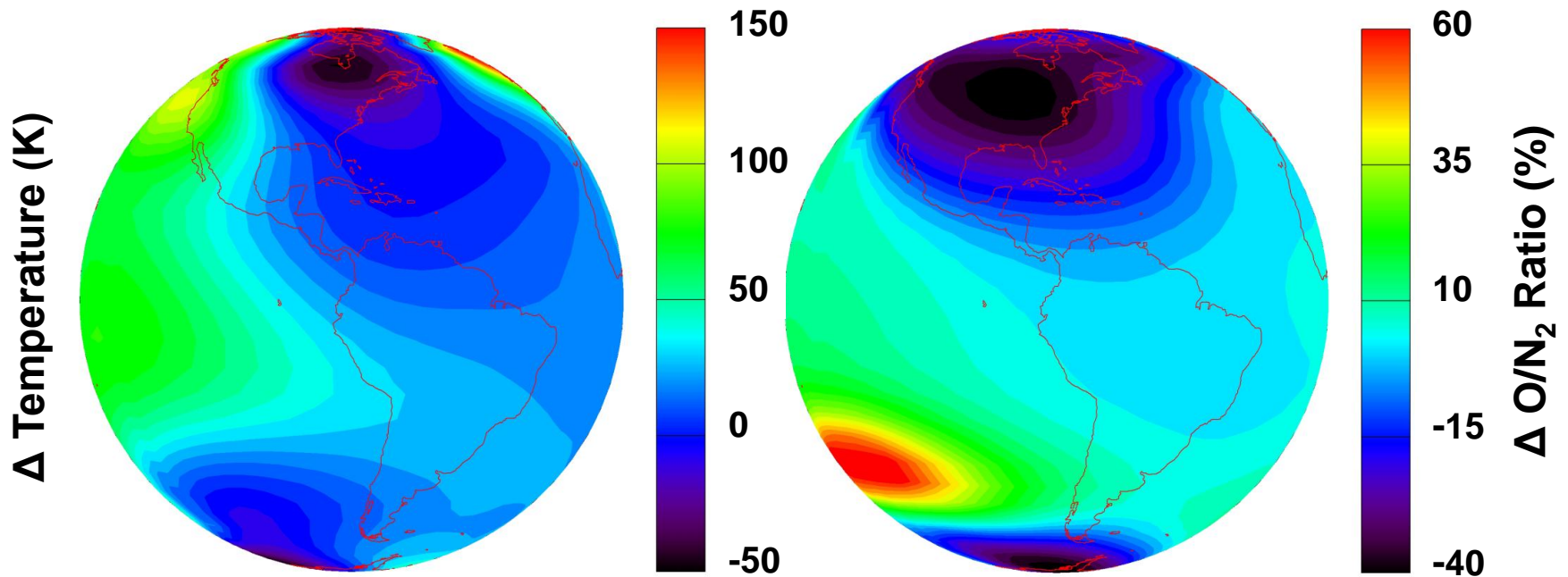


## Thermospheric Neutral Temperatures from MSIS Model



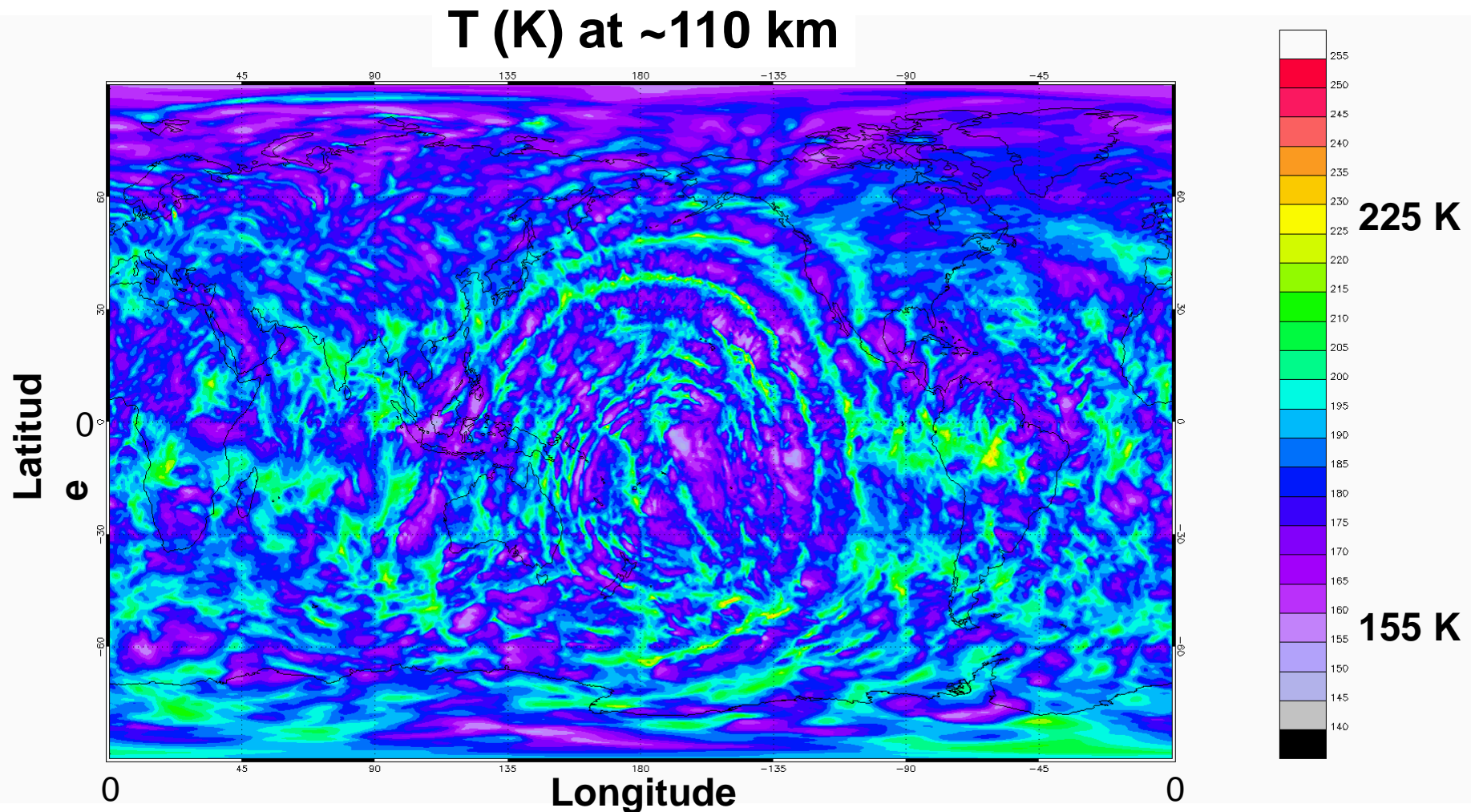
**GOLD simultaneously images key parameters - temperature and composition ( $O/N_2$ ) - in lower thermosphere on the dayside disk**

How do geomagnetic storms impact Earth's space environment?



Modeled changes in upper atmosphere during storm

***GOLD will discover how the upper atmosphere acts as a weather system***



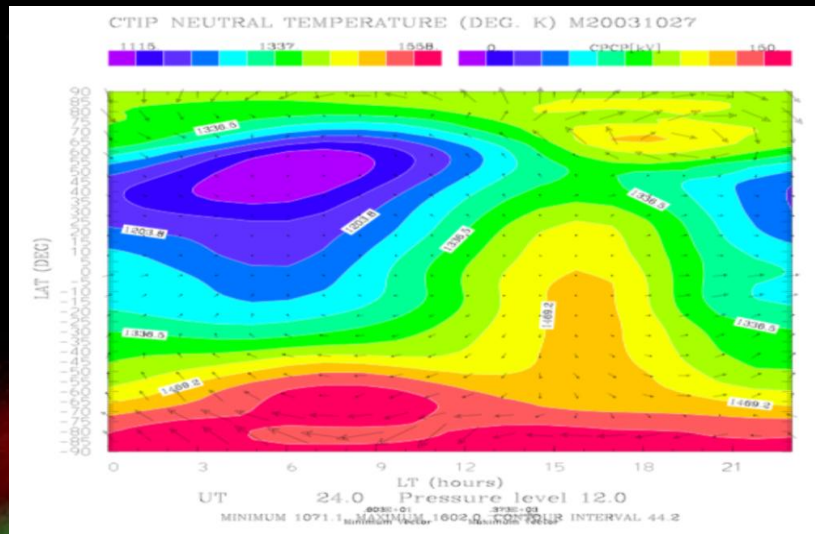
**WACCM Calculation of Gravity Waves at High Resolution  
(0.25° Spatial by 0.1 Scale Height)**



# With and Without Lower Atmosphere:

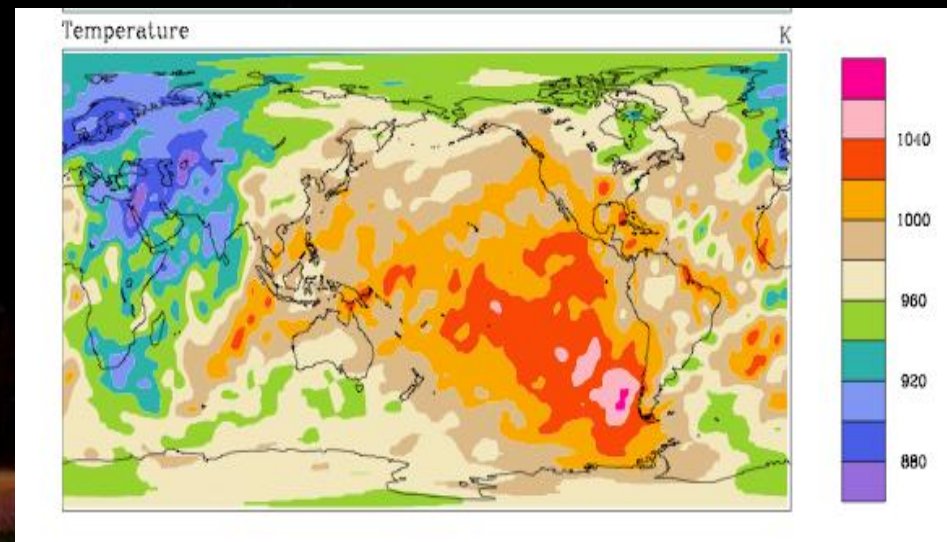
Typical iono-thermosphere model:

- Driven by Solar EUV and Geomagnetic Storms.
- Global maps show little fine structure



Ionosphere-thermosphere model coupled to the lower atmosphere: Global maps show structure relevant to

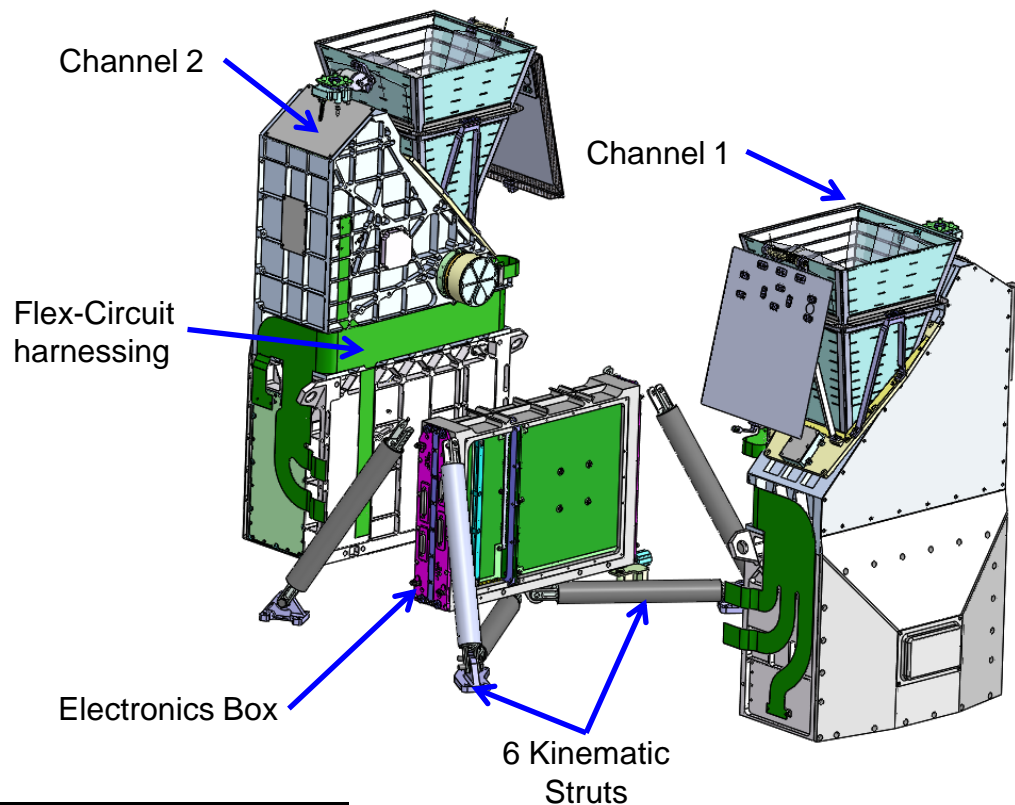
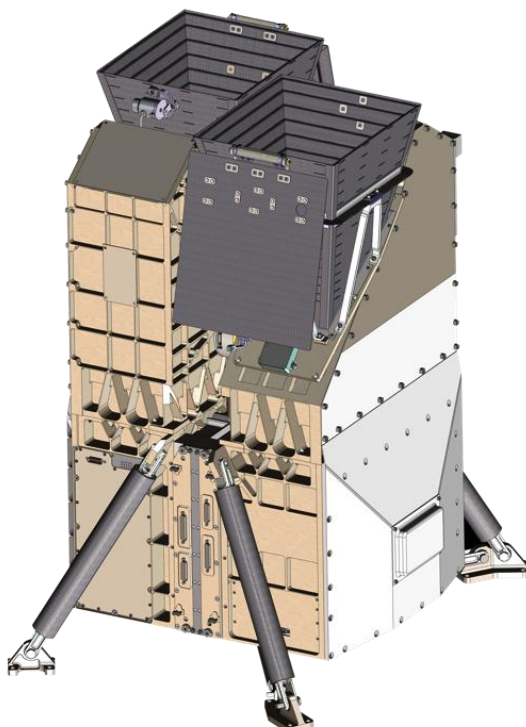
- GPS accuracy and availability
- HF Comm.



The temperature structure from a stand-alone thermosphere ionosphere plasmasphere model (e.g., CTIPe) is similar to the MSIS empirical model. The Whole Atmosphere Model (WAM) drives variability from the chaotic lower atmosphere which introduces a whole spectrum of variability.



- Imaging Spectrograph: Two independent, identical channels imaging the limb and disk, and a single processor packaged in one housing
- Wavelength range: 132 – 160 nm
- Detectors: Microchannel plate, 2-D crossed delay line anode



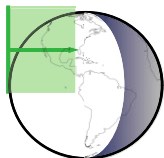
Instrument Summary	
Mass	33 kg (CBE)
Power	51 W (CBE, AVG)
Size	51 × 55 × 69 cm <sup>3</sup>

**Instrument sensitivity is a factor of ~2 better than at CSR due to updated design.**

- Full disk images and limb scans with 30 minute cadence
  - Dayside data products: Disk Temperature, Disk O/N<sub>2</sub>, OI and N<sub>2</sub> emission brightness,  $T_{\text{EXO}}$ ,  $Q_{\text{EUV}}$
  - Nightside products: Disk OI brightness, crest locations,  $N_{\text{max}}$
- Occultation measurements
  - Dayside and nightside products: O<sub>2</sub> density profile

## Dayside Disk Imaging Temperature & O/N<sub>2</sub> Ratio

1



LIMB/HR slit in place  
N hem. day, W to E  
Interrupt for star occultation

## Star Occultation O<sub>2</sub> Density Profile

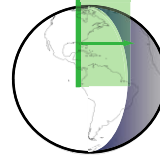
2



OCC slit in place  
OCC

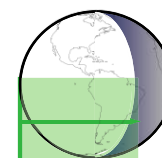
## Dayside Disk Imaging Temperature & O/N<sub>2</sub> Ratio

3a



Return to HR scan location  
LIMB/HR slit in place  
HR scan, N hem. day, W to E

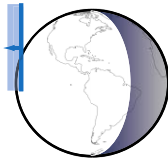
3b



HR scan, S hem. day, W to E

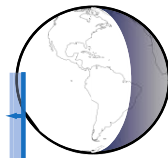
## Dayside Limb Scan N<sub>2</sub> Emission Profile

4a



LIMB/HR slit in place  
NW region, E to W

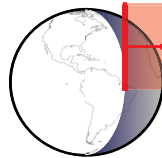
4b



SW region, E to W

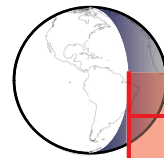
## Nightside Disk Imaging Ionospheric Irregularities

5a



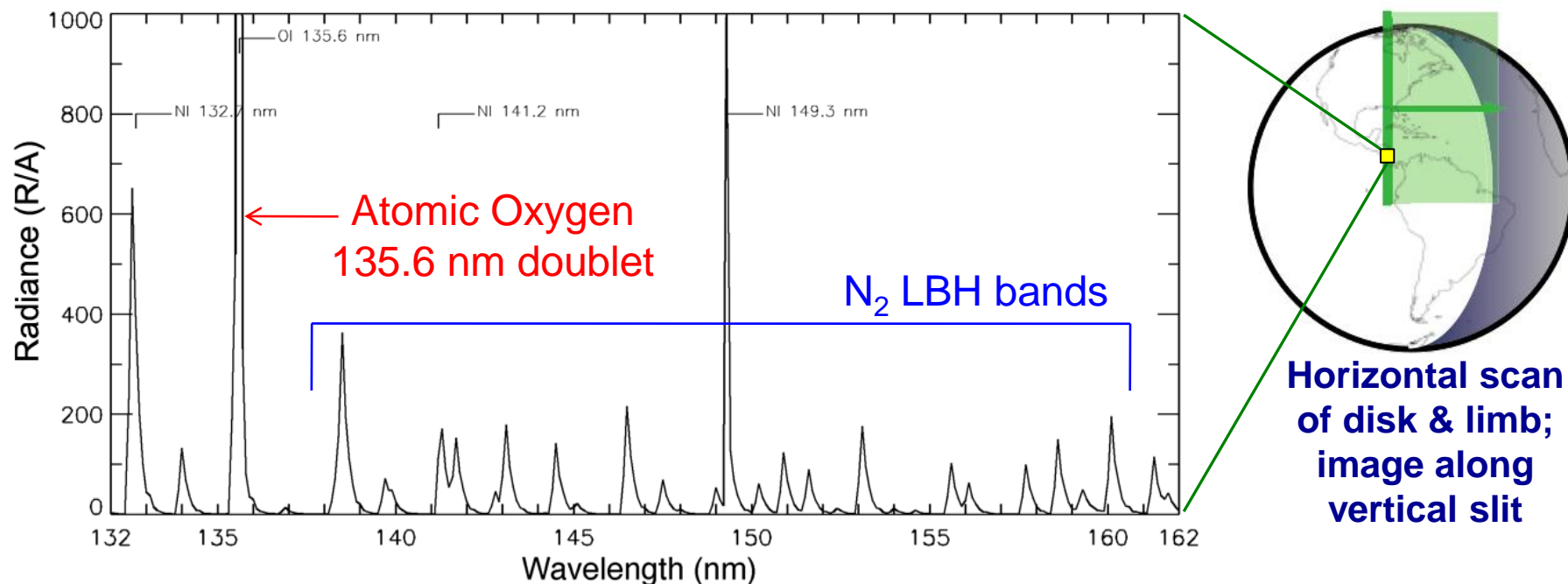
LR slit in place  
N hem., W to E

5b

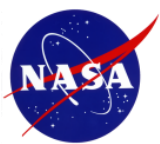


S hem., W to E

## Daytime Far-Ultraviolet Spectrum



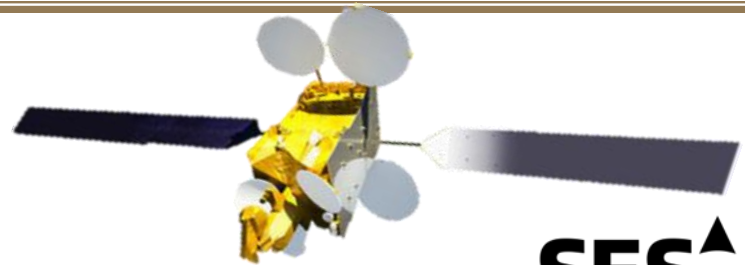
- Temperature obtained on disk from rotational shape of  $N_2$  LBH bands
- O/ $N_2$  composition measured using ratio of 135.6 doublet to LBH bands
- Temperature on limb determined by slope of emission altitude profile
- $O^+$  at night observed using 135.6 recombination emission
- $O_2$  profile on limb from stellar occultations



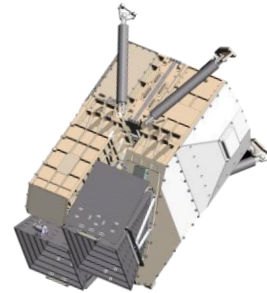
# GOLD Mission Space Segment

**GOLD**

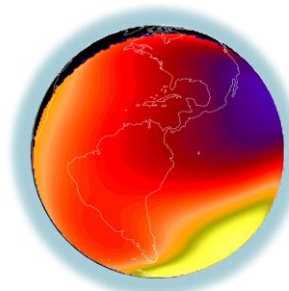
- Host Mission
  - Managed by SES
  - Host Accommodation will be on SES-14
    - GEO commercial communications satellite at 47.5°W, owned and operated by SES
    - Host satellite is a Eurostar 3000 built by Airbus Defence & Space
- GOLD Mission Instrument
  - Hosted Payload is an ultraviolet imager developed by UCF/LASP
  - 6 Mbit/s data down-link
- Data processing at UCF



**SES**<sup>▲</sup>  
Government Solutions

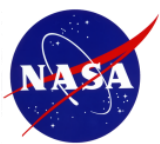


**LASP**



 **UCF**





# **GOLD Mission Summary & Status**

**GOLD**

- **Launches in 2017 for a two-year mission**
- **Unprecedented, simultaneous imaging of composition and temperature**
- **Able to separate changes in time from changes in location**
- **Capability for continuous, real time data availability is inherent to the mission**
- **Provides context for ground-based and LEO measurements**
- **Coincident with ground based and LEO missions, ICON**
- **Mission confirmed on March 5**
- **Spacecraft accommodation contract has been signed**